Note 140
October 1971 and a Two-Parallel-Plate Simulator, I. (Top Plate Removed)
by

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## Abstract

The proximity effect is studied of a perfectly conducting plate on the response of a perfectly conducting sphere illuminated by a monochromatic plane wave or a step-function pulse. Numerical results are obtained for the frequency and time response of the charge densities on the poles of the sphere and of the total current crossing the equator.

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## I. Introduction

In this note we examine the interaction of a plane electromagnetic wave or a step-function pulse with a perfectly conducting sphere within a two-parallel-plate simulator. In particular we focus our attention on the special case where the sphere has a diameter much smaller than the plate separation and is situated close to the bottom plate. We idealize this problem by completely removing the top plate. In a future note we will consider the twoplate interaction problem. The interaction of a test object within a two-parallel-plate simulator with an incident wave has been studied before in both the static and dynamic regimes. The test object has been taken to be a cylinder of finite length (Ref. 1, 2, 3, 4, 5, 6) or infinite length (Ref. 7). Returning to the special problem where the sphere is situated above a perfectly conducting plate and is illuminated by an incident wave, we can use the method of images to arrive at the interaction problem involving two perfectly conducting spheres of equal radius. Studies of the multiple scattering of an incident wave by two spheres exist in the literature (see for example Ref. 8, 9, 10, 11). The quantities of interest in these studies are the scattered field and the backscattering cross section. In this note we are primarily interested in the proximity effect of the plate on the sphere. Thus, in studying this effect we can select some specific quantities the knowledge of which requires a relatively simple formulation. These quantities are the charge densities on the north and south poles and the total current crossing the equator. The knowledge of the charge densities at these points furnishes information about the local interaction in the immediate vicinity of the plates, whereas the knowledge of the current provides an overall picture of the effect of the plates on the interaction problem. Our formulation consists of deriving an integral equation for the $\phi$ independent component of the current density. The angle $\phi$ is the azimuthal angle in a spherical coordinate system with the $z$ axis piercing through the south and north poles. The quantities of interest previously mentioned, that is, the charge densities on the north and south poles and the total current can be derived from a knowledge of the $\phi$ independent current density.

In section II we give a formulation for the general interaction problem involving a perfectly conducting body within a two-parallel-plate simulator.

This is done by deriving an integral equation for the current density on the conducting body. Next we specialize to the sphere case and we present an integral equation for the $\phi$ independent current density. Finally, we remove the top plate and focus our attention on the one-plate problem. We solve the integral equation in the frequency domain numerically on a computer, and we obtain the time response to an incident step-function pulse by numerical Fourier inversion. In section III, we discuss the interrelationship between high and low-frequency behavior to early and late-time behavior. In section IV we discuss the plots of the quantities of interest in the frequency and time domain and ascertain the importance of the proximity effect of the plate.
II. Formulation of the Problem

In general one has to solve the problem of the scattering of a plane monochromatic wave by a conducting body positioned between two parallel conducting plates of infinite extent (Fig. 1). An integral equation approach for the current induced on the conducting body is feasible through the use of a dyadic Green's function satisfying appropriate boundary conditions on the plates. We start from the pair of equations

$$
\begin{array}{r}
\nabla \times \nabla \times \underline{H}-k^{2} \underline{H}=\nabla \times \underline{I}_{i} \\
\nabla \times \nabla \times \underline{\underline{G}}-k^{2} \underline{\underline{G}}=\underline{\underline{u}} \delta\left(\underline{r}-\underline{\underline{r}}^{\prime}\right) \tag{2}
\end{array}
$$

where $J_{i}$ is the current source for the incident field and $\underset{=}{\underline{=}}$ is the unit dyadic. The Green's function $\underline{\underline{G}}$ is assumed to satisfy

$$
\begin{equation*}
\underline{\underline{n}} \times \operatorname{curl} \underline{\underline{G}}=0 \quad \text { on the plates } \tag{3}
\end{equation*}
$$

If we apply the vector Green's theorem for the pair (1) and (2) we obtain

$$
\begin{align*}
& =\oint_{S}\left[\left(\underline{n^{\prime}} \times \underline{H}\left(\underline{r^{\prime}}\right)\right) \cdot \operatorname{curI} \underline{\underline{G}}(\underline{\underline{r}} ; \underline{\underline{r}})+\left(\underline{n}^{\prime} \times \operatorname{curl} \underline{\underline{H}}(\underline{\underline{r}})\right) \cdot \underline{\underline{G}}\left(\underline{r^{\prime}} ; \underline{\underline{r}}\right)\right] d S^{\prime} \tag{4}
\end{align*}
$$

where $S=S_{1}+S_{2}+S_{3}+S_{4}+S_{B}$ as indicated in figure 1 . We first -note that ( $\underline{n}^{\prime} \times \underline{H}$ ) $\operatorname{curl} \underline{\underline{G}}=-\underline{H} \cdot\left(\underline{n}^{\prime} \times \operatorname{curl} \underline{\underline{G}}\right.$ ). We also recall that $\underline{n}^{\prime} \times \operatorname{curl} \underline{H}=0$ on a perfect conductor. Using these facts and (1), (2) and (3), (4) can be rewritten as

$$
\begin{align*}
\underline{H}(\underline{r})-\underline{H}_{i}(\underline{r})= & \int_{S_{3}+S_{4}}\left[\left(\underline{n^{\prime}} \times \underline{H}\right) \cdot \operatorname{curl} \underline{\underline{G}}+\left(\underline{n^{\prime}} \times \operatorname{curl} \underline{\underline{H}}\right) \cdot \underline{\underline{G}}\right] d S^{\prime} \\
& +\int_{S_{B}}\left(\underline{\underline{n}}^{\prime} \times \underline{H}\right) \cdot \operatorname{curl} \underline{\underline{G} S^{\prime}} \tag{5}
\end{align*}
$$

where

$$
\underline{H}_{i}(\underline{x})=\int_{V_{0}} \nabla^{\prime} \times \underline{J}_{i} \cdot \underline{G}^{G}
$$

The integral over $S_{3}+S_{4}$ vanishes due to the radiation condition and (5) reduces to

$$
\begin{equation*}
\underline{H}(\underline{r})=\underline{H}_{i}(\underline{r})+\int_{S_{B}}\left[\underline{\underline{r}}^{\prime} \times \underline{H}\left(\underline{\underline{r}}^{\prime}\right)\right] \cdot \operatorname{cur} 1 \underline{\underline{G}}\left(\underline{\underline{r}}^{\prime} ; \underline{r}\right) \mathrm{dS}^{\prime} \tag{6}
\end{equation*}
$$

We can now remove the source $\mathrm{J}_{\mathrm{i}}$ to infinity to obtain a monochromatic plane wave as our incident field with the $H_{i}$ field parallel to the plates. Equation (6) can be converted to an integral equation for the current density $\underline{K}(\underline{r})=\underline{n} \times \underline{H}(\underline{r})$ in the usual manner by letting $\underline{r}$ approach the surface of the body.

In this note we are interested in some field quantities which do not require a detailed knowledge of $\underline{K}(\underline{r})$. Our conducting body is a sphere of radius b (Fig. 2); we want to calculate the electric field or equivalently the charge density on the north and south pole and also the total current. These quantities only depend on the $\phi$ independent component of the current density $\underline{K}(\underline{r})$ and consequently we will seek a formulation for this component. $\underline{K}(\theta)$ is directed along the $\theta$ direction and it only depends on the $\phi$ independent component of the incident magnetic field $H_{i}=-H_{0} e^{i k x} e_{y}$. The $\phi$ independent magnetic field has a $\phi$ component only and satisfies the following equation

$$
\begin{equation*}
\frac{1}{r} \frac{\partial^{2}}{\partial r^{2}}\left(r H_{\phi}\right)+\frac{1}{r^{2}} \frac{\partial}{\partial \theta}\left[\frac{1}{\sin \theta} \frac{\partial}{\partial \theta}\left(\sin \theta H_{\phi}\right)\right]+k^{2} H_{\phi}=0 \tag{7}
\end{equation*}
$$

or

$$
\left(\mathrm{L}+\mathrm{k}^{2}\right) \mathrm{H}_{\phi}=0
$$

where the meaning of $L$ is obvious. We first notice that

$$
\begin{equation*}
L=\nabla^{2}-\frac{1}{r^{2} \sin ^{2} \theta}-\frac{1}{r^{2} \sin ^{2} \theta} \frac{\partial^{2}}{\partial \phi^{2}} \tag{8}
\end{equation*}
$$

where $\nabla^{2}$ is the Laplacian operator. In view of (8) we understand that $u=H_{\phi} e^{i \phi}$ satisfies the scalar Helmholtz equation

$$
\begin{equation*}
\left(\nabla^{2}+k^{2}\right) u=0 \tag{9}
\end{equation*}
$$

Next we introduce a Green's function $G_{1}$ satisfying

$$
\begin{equation*}
\left(\nabla^{2}+k^{2}\right) G_{1}=-\delta\left(\underline{r}-\underline{\underline{r}}^{\prime}\right) \tag{10}
\end{equation*}
$$

with a boundary condition

$$
\begin{equation*}
\frac{\partial G_{1}}{\partial z}=0 \quad \text { on the plates } \tag{11}
\end{equation*}
$$

Applying Green's theorem for the pair (9) and (10) we arrive as before at the following equation

$$
\begin{equation*}
u(\underline{r})=u_{i}(\underline{r})+\int_{S_{1}+S_{2}}\left(u \frac{\partial G_{1}}{\partial z^{\prime}}-G_{1} \frac{\partial u}{\partial z^{\prime}}\right) d S^{\prime}+\int_{S_{B}}\left(u \frac{\partial G_{1}}{\partial r^{\prime}}-G_{1} \frac{\partial u}{\partial r^{\prime}}\right) d S^{\prime} \tag{12}
\end{equation*}
$$

From Maxwell's equation $-i \omega \varepsilon_{0} E=\nabla \times \underline{H}$, the $\phi$ independent components $E_{\rho}$ and $H_{\phi}$ (where $\rho$ is the polar radius in a suitable cylindrical coordinate system) satisfy $i \omega \varepsilon_{0} E_{\rho}=\left(\partial H_{\phi} / \partial z\right)$ and consequently

$$
\begin{equation*}
\frac{\partial \mathrm{H}_{\phi}}{\partial z}=0 \quad \text { on the plates } \tag{13}
\end{equation*}
$$

If we express $-i \omega \varepsilon_{0} E=\nabla \times \underline{H}$ in spherical coordinates we find $i \omega \varepsilon_{0} r E_{\theta}=(\partial / \partial r)\left(r H_{\phi}\right)$, i.e.,

$$
\begin{equation*}
\frac{\partial}{\partial r}\left(\mathrm{rH}_{\phi}\right)=0 \quad \text { on the sphere } \tag{14}
\end{equation*}
$$

In view of the boundary conditions satisfied by the field $\underline{H}$ and $G_{1}$ (12) can be rewritten as

$$
\begin{equation*}
u(\underline{r})=u_{i}(\underline{r})+\left.\int_{S_{B}} u\left(\underline{r}^{\prime}\right) \frac{1}{r^{\prime}} \frac{\partial}{\partial r^{\prime}}\left(r^{\prime} G_{1}\right)\right|_{r^{\prime}=b} \tag{15}
\end{equation*}
$$

where $d S^{\prime}=b^{2} \sin \theta^{\prime} d \theta^{\prime} d \phi^{\prime}$. The $\phi$ independent $\phi$ component of the incident magnetic field $-H_{o-y} e^{i k x}$ is

$$
\frac{1}{2 \pi}\left(-H_{0}\right) \int_{0}^{2 \pi}\left(e_{y} \cdot e_{\phi}\right) e^{i k r \sin \theta \cos \phi} d \phi=-i H_{0} J_{1}(k r \sin \theta) .
$$

Multiplying now both sides of (15) by ( $1 / 2 \pi$ ) $e^{-i \phi}$ and integrating from 0 to $2 \pi$ with respect to $\phi$ we obtain

$$
\begin{align*}
& H_{\phi}(r, \theta)=-i H_{0} J_{1}(k r \sin \theta) \\
& +\frac{1}{2 \pi} \int_{0}^{\pi} d \theta^{\prime} H_{\phi}\left(b, \theta^{\prime}\right) b \sin \theta^{\prime}\left\{\frac{\partial}{\partial r^{\prime}}\left[r^{\prime} \int_{0}^{2 \pi} d \int_{0}^{2 \pi} d \phi^{\prime} e^{i\left(\phi^{\prime}-\phi\right)} G_{1}\left(r^{\prime}, \theta^{\prime}, \phi^{\prime} ; r, \theta, \phi\right)\right]\right\}_{r^{\prime}=b}^{(16} \tag{16}
\end{align*}
$$

The Green's function $G_{I}$ is an even function of $\phi$ ' - $\phi$ and consequently $\int_{0}^{2 \pi} d \phi \int_{0}^{2 \pi} d \phi^{\prime} G_{1}\left(\left|\phi-\phi^{\prime}\right|\right) e^{i\left(\phi^{\prime}-\phi\right)}=2 \pi \int_{0}^{2 \pi} d \psi G_{1}(\psi) e^{i \psi \psi}=2 \pi \int_{0}^{2 \pi} d \psi \cos \psi G_{1}(\psi)$.

Equation (16) now reduces to

$$
\begin{equation*}
H_{\phi}(r, \theta)=-i H_{0} J_{1}(k r \sin \theta)+\int_{0}^{\pi} H_{\phi}\left(b, \theta^{\prime}\right)\left[\frac{\partial}{\partial r^{\prime}}\left(r^{\prime} G\right)\right]_{r^{\prime}=b} b \sin \theta^{\prime} d \theta^{\prime} \tag{17}
\end{equation*}
$$

where

$$
\begin{equation*}
G=\int_{0}^{2 \pi} d \psi \cos \psi G_{1}\left(r, \theta, r^{\prime}, \theta^{\prime}, \psi\right) \tag{18}
\end{equation*}
$$

We now bring $r$ onto the surface of the sphere and (17) is converted in the usual manner into an integral equation for $K_{\theta}=-H_{\phi}(b, \theta)$.

$$
\begin{equation*}
\frac{1}{2} K_{\theta}(\theta)+\int_{0}^{\pi} M\left(\theta ; \theta^{\prime}\right) K_{\theta}\left(\theta^{\prime}\right) d \theta^{\prime}=i J_{1}(k b \sin \theta) H_{0} \tag{I9}
\end{equation*}
$$

with

$$
\begin{equation*}
M\left(\theta ; \theta^{\prime}\right)=-b \sin \theta^{\prime}\left[\frac{\partial}{\partial r^{\prime}}\left(r^{\prime} G_{1}\right)\right]_{r^{\prime}=b} \tag{20}
\end{equation*}
$$

and $G$ is given by (18) with $r=b$. The Green's function $G_{1}$ can be computed by the method of images and is given by

$$
\begin{equation*}
G_{1}\left(b, \theta, r^{\prime}, \theta^{\prime}, \psi\right)=\sum_{m=-\infty}^{\infty} \frac{e^{i k R_{1 m}}}{4 \pi R_{1 m}}+\sum_{m=0}^{\infty}\left(\frac{e^{i k R_{2 m}}}{4 \pi R_{2 m}}+\frac{e^{i k R_{3 m}}}{4 \pi R_{3 m}}\right) \tag{21}
\end{equation*}
$$

with

$$
\begin{align*}
& \mathrm{R}_{1 \mathrm{~m}}=\left[\left(2 \mathrm{~ms}+z^{\prime}-z\right)^{2}+\rho^{2}+\rho^{\prime 2}-2 \rho \rho^{\prime} \cos \psi\right]^{\frac{1}{2}} \\
& \mathrm{R}_{2 \mathrm{~m}}=\left[\left(2 \mathrm{~ms}+2 \mathrm{~h}_{1}-z^{\prime}-z\right)^{2}+\rho^{2}+\rho^{\prime 2}-2 \rho \rho^{\prime} \cos \psi\right]^{\frac{1}{2}}  \tag{22}\\
& \mathrm{R}_{3 \mathrm{~m}}=\left[\left(2 \mathrm{~ms}+2 \mathrm{~h}_{2}+z^{\prime}+z\right)^{2}+\rho^{2}+\rho^{\prime 2}-2 \rho \rho^{\prime} \cos \psi\right]^{\frac{1}{2}}
\end{align*}
$$

where $z=b \cos \theta, z^{\prime}=r^{\prime} \cos \theta^{\prime}, \rho=b \sin \theta, \rho^{\prime}=b \sin \theta^{\prime}, h_{1}+h_{2}=s$. The current $I(\theta)$ is given by

$$
I(\theta)=\int_{0}^{2 \pi} K_{\theta}(\theta, \phi) b \sin \theta d \phi
$$

If we expand $K_{\theta}(\theta, \phi)$ in a Fourier series, then we see that $I(\theta)$ only depends on the $\phi$ independent component of $\mathrm{K}_{\theta}(\theta, \phi)$. Thus,

$$
\begin{equation*}
I(\theta)=2 \pi b \sin \theta K_{\theta}(\theta) \tag{23}
\end{equation*}
$$

The current density $\sigma$ at a given point $(\theta, \phi)$ does depend on $\phi$. For $\theta=0, \pi$ we should obtain the same value for $\sigma$ no matter what $\phi$ is, and $\sigma$ cannot depend on $\phi$ at these two points. Thus, $\sigma$ at $\theta=0, \pi$ coincides with the $\phi$ independent component of the charge density. It can be obtained through the continuity equation

$$
\begin{equation*}
i \omega \sigma=\frac{1}{b \sin \theta} \frac{\partial}{\partial \theta}\left(\sin \theta K_{\theta}\right) \tag{24}
\end{equation*}
$$

where $K_{\theta}$ and $\sigma$ are the $\phi$ independent current density and charge density respectively. First we calculate $\sigma$ at $\theta=0$ 。 We rewrite (24) as

$$
i \omega \sigma=\frac{1}{b} \frac{\partial K_{\theta}}{\partial \theta}+\frac{\cos \theta}{\sin \theta} K_{\theta}
$$

As $\theta \rightarrow 0$,

$$
\left.\frac{\cos \theta}{\sin \theta} K_{\theta} \rightarrow \frac{\partial K_{\theta}}{\partial \theta}\right|_{\theta=0}
$$

and (24) becomes

$$
\begin{equation*}
\sigma=-\left.\frac{2 i}{\omega b} \frac{\partial K_{\theta}}{\partial \theta}\right|_{\theta=0} \tag{25}
\end{equation*}
$$

$K_{\theta}$ is zero at $\theta=0$, therefore,

$$
\frac{\partial K_{\theta}}{\partial \theta}=\left.\frac{K_{\theta}}{\theta}\right|_{\theta=0}
$$

or

$$
\begin{equation*}
\sigma=-\left.\frac{2 i}{\omega b} \frac{K_{\theta}}{\theta}\right|_{\theta=0} \tag{26}
\end{equation*}
$$

From the numerical computation point of view (25) or (26) cannot give accurate results unless (19) is numerically integrated using a large number of zones. We can improve the accuracy by calculating $\sigma$ as an integral over $k_{\theta}$. If we combine (19) and (25) we obtain

$$
\begin{equation*}
\sigma=2 \frac{H_{0}}{c}+\int_{0}^{\pi} N\left(\theta^{\prime}\right) K_{\theta}\left(\theta^{\prime}\right) d \theta^{\prime} \tag{27}
\end{equation*}
$$

where

$$
\begin{equation*}
N\left(\theta^{\prime}\right)=-\frac{4 i}{\omega b}\left[\frac{\partial}{\partial \theta} M\left(\theta, \theta^{\prime}\right)\right]_{\theta=0} \tag{28}
\end{equation*}
$$

Noticing that $H_{0} / E_{0}=\left(\varepsilon_{0} / \mu_{0}\right)^{\frac{1}{2}}$, i.e. $H_{0} / c=\varepsilon_{0} E_{0}$ (27) becomes

$$
\begin{equation*}
\sigma(\theta=0)=2 \varepsilon_{0} E_{0}+\int_{0}^{\pi} N\left(\theta^{\prime}\right) K_{\theta}\left(\theta^{\prime}\right) d \theta^{\prime} \tag{29}
\end{equation*}
$$

For a given accuracy of $\mathrm{K}_{\theta}$ the charge density $\sigma(\theta=0)$ calculated through (29) is more accurate than $\sigma$ given by (26). To calculate $\sigma(\theta=\pi)$ we use (29) with the $\theta$ dependence of $K_{\theta}$ reversed. The result is equal to $-\sigma(\theta=\pi)$.

In this note we will consider the case of one plate only (Fig. 3), and in a future note we will study the two-plate problem. The formulation is still the same except that $G_{1}$ is considerably simplified. Instead of (21) we now have

$$
\begin{equation*}
G_{1}\left(b, \theta, r^{\prime}, \theta^{\prime}, \psi\right)=\frac{e^{i k R_{10}}}{4 \pi R_{10}}+\frac{e^{i k R_{30}}}{4 \pi R_{30}} \tag{30}
\end{equation*}
$$

where

$$
\begin{align*}
R_{10} & =\left[\left(z-z^{\prime}\right)^{2}+\rho^{2}+\rho^{\prime 2}-2 \rho \rho^{\prime} \cos \psi\right]^{\frac{1}{2}} \\
R_{30} & =\left[\left(2 h_{2}+z^{\prime}+z\right)^{2}+\rho^{2}+\rho^{\prime 2}-2 \rho \rho^{\prime} \cos \psi\right]^{\frac{1}{2}}  \tag{31}\\
z & =b \cos \theta, \quad z^{\prime}=r^{\prime} \cos \theta^{\prime}, \quad \rho=b \sin \theta, \quad \rho^{\prime}=r^{\prime} \sin \theta^{\prime}
\end{align*}
$$

When $r^{\prime}=b, \theta=\theta^{\prime}$ then $R_{10}=0$, and the Kernel in (17) is singular. This is an integrable singularity and in the Appendix we show how to treat this singularity in order to perform the integration numerically on a computer.
III. Low, High-Frequency and Late, Early-Time Behavior

For a plane wave of low frequency the wavelength is much larger than the radius of the sphere and we have a quasi-static situation. Consider now a step pulse illuminating the sphere. At late times the pulse has well enveloped the sphere and the induced charges have again settled to a quasi-static state. Thus, we expect the current and charge densities to asymptotically reach their static values at late times. This value is zero for the current. For a plane wave of high frequency the wavelength is much smaller than the radius of the sphere and the interaction has a local character. The current density and charge density are $2 \underline{n} \times \underline{H}^{i n c}$ and $2 \underline{n} \cdot \underline{E}^{i n c}$ in the illuminated region and zero in the shadow region. We are interested in the values of the charge density at $\theta=0$ and $\pi$, i.e., right on the shadow boundary. Fock (Ref. 10) and others (see for example Ref. 11) have worked out the values of the current density in the transition region. They have found that in this region, $\underline{K}(\theta, \phi)=\underline{K}^{\text {inc }}(\theta, \phi) G(\xi)$ where $\xi$ is negative in the illuminated region and positive in the shadow region. At $\xi=0$, i.e., on the shadow boundary $G(\xi)=1.3990$ whereas for $\xi$ large and negative $(|\xi|>3) G(\xi) \approx 2$ and for $\xi$ large and positive $(\xi>5) G(\xi) \approx 0$. In view of (26) the same considerations are valid for the charge density at $\theta=0$ and aiso $\theta=\pi$. Thus, the value of $\sigma(\theta=0, \pi)$ is $1.3990 \varepsilon_{0} E_{0}$ as $k b \rightarrow \infty$. Consider now a step pulse incident upon the sphere. As the front comes in contact with the sphere the interaction has local character. Thus, the charge density at the poles (at $t=0+$ ) should assume the high frequency value $1.3990 \varepsilon_{0} E_{0}$. Notice that the presence of the plates does not change this value. The total current varies as $k^{-\frac{1}{2}}$ and goes to zero for either $k b \rightarrow \infty$ or $t=0+$. This can be seen by examining (19).

## IV. Discussion of Plots

We present plots of the charge densities at the poles and the equator current. The pole closest to the ground plate is designated as B (bottom) and the other pole as $T$ (top). We normalize the charge densities either dividing them by $\varepsilon_{0} E_{o}$ or by the appropriate free space value. The equator current is denoted as I and the normalizing factor is either $b H_{o}$ or the free space value. Thus, we give the following plots

Figure $7 \quad \sigma_{B}(t) / \varepsilon_{0} E_{0}$ versus $c t / b$
Figure $8 \quad\left|\sigma_{B}(\omega)\right| / \varepsilon_{o} E_{0}$ versus kb
Figure $9,10\left|\sigma_{B}(\omega)\right| / \varepsilon_{0} E_{0}$ versus $k b$, expanded scale
Figure $11 \sigma_{T}(t) / \varepsilon_{0} E_{0}$ versus $c t / b$
Figure $12\left|\sigma_{T}(\omega)\right| / \varepsilon_{0} E_{0}$ versus $k b$
Figure $13 \quad I(t) / b H_{o}$ versus $c t / b$
Figure $14|I(\omega)| / \mathrm{bH}_{\mathrm{o}}$ versus kb
Figure 15 max $\sigma_{B}(t) / \varepsilon_{0} E_{o}$ versus $\delta / b$ where $\max \sigma_{B}(t)$ is the maximum value attained by $\sigma_{B}(t)$ in the time domain
Figure $16 C_{m}(B)$ versus $\delta / b$ where $C_{m}(B)$ is the ratio of the maximum $\sigma_{B}(t)$ for a given $\delta / b$ to the maximum $\sigma_{B}(t)$ for $\delta / b=\infty$ (free space)
Figure $17 \max \left|\sigma_{B}(\omega)\right| / \varepsilon_{0} E_{o}$ versus $\delta / b$ where $\max \left|\sigma_{B}(\omega)\right|$ is the maximum value of $\left|\sigma_{B}(\omega)\right|$ corresponding to the first resonance
Figure $18 \quad C_{r}(B)$ versus $\delta / b$ where $C_{r}(B)$ is the ratio of the maximum $\left|\sigma_{B}(\omega)\right|$ for a given $\delta / b$ to the maximum $\left|\sigma_{B}(\omega)\right|$ for $\delta / b=\infty$ (free space)
Figure $19 \sigma_{B}(\omega=0) / \varepsilon_{0} E_{0}$ versus $\delta / b$ where $\sigma_{B}(\omega=0)$ is the static value at $B$
Figure $20 \quad C_{S}(B)$ versus $\delta / b$ where $C_{S}(B)$ is the ratio of the static value of $\sigma_{B}$ for a given $\delta / b$ to the static value of $\sigma_{B}$ for $\delta / b=\infty$ (free space)
Figure 21 through Figure 26 are the same as 15 through 20 except they refer to the charge density at $T$
Figure $27 \max |I(\omega)| / \mathrm{bH}_{0}$ versus $\delta / b$ where $\max |I(\omega)|$ is the value of $|I(\omega)|$ at the first resonance

Figure $28 C_{r}(I)$ versus $\delta / b$ where $C_{r}(I)$ is the ratio of the value of $I$ for $a$ given $\delta / b$ at the first resonance to the value of $I$ for $\delta / b=\infty$ (free space) at the corresponding first resonance.

No plots of max $I(t)$ versus $\delta / b$ or $C_{m}(I)$ versus $\delta / b$ are given because the percentage differences between the maximum values of $I$ for various $\delta / b$ are small ( $\leq 7.5 \%$ )

The zero time for the charge densities is taken as the moment of the arrival of the step pulse wavefront at the center of the sphere and the zero time for the current is the moment of contact of the wavefront with the sphere. The effect of the ground plate or equivalently of the image sphere occurs at a retarded time depending on the quantity of interest and the gap size. The computation of the retardation times are given in Appendix B. For the pole $B$ the ray that travels the least time is the one that suffers a specular reflection at a point on the surface of the image sphere (Fig. 4 and 5). In Appendix $C$ we give a proof that any deviation from the specular reflection path results in a longer retarded time. A direct specular reflection path Is not feasible for either the points on the equator or $T$. For the current the least time path consists of a specular reflection off the image sphere such that the reflected ray will be tangent to the real sphere plus an arc path along the surface of the latter sphere (Fig. 6). To obtain the minimum time path for point $T$, at the top of the sphere, we should add an arc of $90^{\circ}$ to the current path.

In the frequency domain the presence of the ground plate can have a pronounced effect on the amplitude of the charge density $\left|\sigma_{B}(\omega)\right|$ but a less pronounced of the first resonant frequency. The corresponding effects on $\left|\sigma_{T}(\omega)\right|$ and $|I /(\omega)|$ are much less pronounced for the amplitude but comparable for the first resonant frequency; however, the presence of the ground plane can cause some additional secondary resonances which are not present in the free space case. These resonances appear to be predominant for $\left|\sigma_{T}(\omega)\right|$ and $|I(\omega)|$ onIy. The reason is that additional resonances for $\left|\sigma_{B}(\omega)\right|$ due to the proximity effect can only occur for large $k b$ at which the gap can be comparable to the wavelength and at large $k b$ the amplitude tends to settle down to the high frequency limit which is independent of the presence of the ground plate, i.e., the resonances would appear as minor perturbations. Finally, we would like to make two remarks with respect to the time response of the sphere. First, the time at which the quantity of interest (charge density or current) reaches its
peak value has only a slight dependence on the gap size. Second, the time interval from the peak to a value which has a percentage difference from the final (static) value less than $4 \%$ is short compared to a thin cylindrical structure (see for example Ref. 2) and is given by $\Delta t \simeq 10 b / c$. This last estimate refers to the charge densities since the current settles down to zero for large ct/b.

## Appendix A

The kernel $M\left(\theta ; \theta^{\prime}\right)$ given by (20) can be rewritten as

$$
\begin{equation*}
M\left(\theta ; \theta^{\prime}\right)=\int_{0}^{2 \pi} d \psi \cos \psi M_{1}\left(\theta, \theta^{\prime}, \psi\right), \tag{A-1}
\end{equation*}
$$

where

$$
\begin{equation*}
M_{1}\left(\theta, \theta^{\prime}, \psi\right)=-b \sin \theta^{\prime}\left[G_{10}\left(\frac{1}{2}+\frac{1}{2} i k R_{10}\right)+G_{30}\left(1+i k b \frac{B_{30}}{R_{30}}-\frac{b B_{30}}{R_{30}^{2}}\right)\right] \tag{A-2}
\end{equation*}
$$

and

$$
G_{10}=\frac{e^{i k R_{10}}}{4 \pi R_{10}}, \quad G_{30}=\frac{e^{i k R_{30}}}{4 \pi R_{30}}
$$

$B_{30}=b\left[2\left(h_{2} / b\right) \cos \theta^{\prime}+I+\cos \theta \cos \theta^{\prime}-\sin \theta \sin \theta^{\prime} \cos \psi\right]$,
$R_{10}=\sqrt{2} b\left(1-\cos \theta \cos \theta^{\prime}-\sin \theta \sin \theta^{2} \cos \psi\right)^{\frac{1}{2}}$,
$R_{30}=b\left[\left(2\left(h_{2} / b\right)+\cos \theta+\cos \theta^{\prime}\right)^{2}+\sin ^{2} \theta+\sin ^{2} \theta^{\prime}-2 \sin \theta \sin \theta^{\prime} \cos \psi\right]^{\frac{1}{2}}$.

We understand that the only term singular at $\theta=\theta^{\prime}, \psi=0$ is
$-1 / 2 b \sin \theta^{\prime} G_{10}=-1 / 2 b \sin \theta^{\prime}\left(e^{1 k R_{10}} / 4 \pi R_{10}\right)$. We treat this term as follows. First we define

$$
\begin{equation*}
M_{1}\left(\theta, \theta^{\prime}, \psi\right)=-\frac{1}{2} b \sin \theta^{\prime} G_{10}+M_{2}\left(\theta, \theta^{\prime}, \psi\right) \tag{A-4}
\end{equation*}
$$

and rewrite (19) as

$$
\begin{aligned}
& \frac{1}{2} K_{\theta}(\theta)-\frac{1}{2} b \int_{0}^{\pi}\left[K_{\theta}\left(\theta^{\prime}\right)-K_{\theta}(\theta)\right]\left\{\int_{0}^{2 \pi} d \psi \cos \psi G_{10}\right\} \sin \theta^{\prime} d \theta^{\prime} \\
& -\frac{1}{2} b K_{\theta}(\theta) \int_{0}^{\pi} d \theta^{\prime} \sin \theta^{\prime} \int_{0}^{2 \pi} d \psi \cos \psi G_{10}+\int_{0}^{\pi} K_{\theta}\left(\theta^{\prime}\right) \int_{0}^{2 \pi} d \psi \cos \psi M_{2}\left(\theta, \theta^{\prime}, \psi\right) \\
& =i J_{1}(k b \sin \theta) H_{0}
\end{aligned}
$$

Thus, we have to evaluate

$$
\begin{equation*}
I=\int_{0}^{\pi} d \theta^{\prime} \sin \theta^{\prime} \int_{0}^{2 \pi} d \psi \cos \psi G_{10} \tag{A-6}
\end{equation*}
$$

separately. Define

$$
\begin{equation*}
I_{1}=\int_{0}^{2 \pi} d \psi \cos \psi G_{10}=2 \int_{0}^{\pi} d \psi \cos \psi \frac{e^{i k R_{10}} 10}{4 \pi R_{10}} \tag{A-7}
\end{equation*}
$$

and

$$
\begin{align*}
& A=\left(1-\cos \theta \cos \theta^{\prime}\right)^{\frac{1}{2}} \\
& B=\frac{\sin \theta \sin \theta^{\prime}}{1-\cos \theta \cos \theta^{\prime}} \tag{A-8}
\end{align*}
$$

Thus,

$$
\begin{equation*}
I_{1}=\frac{1}{2(2)^{\frac{1}{2}} \pi A} \int_{0}^{\pi} d \psi \frac{\cos \psi e^{i k R_{10}}}{(1-B \cos \psi)^{\frac{1}{2}}}=\frac{1}{2(2)^{\frac{1}{2}} \pi A} I_{2} \tag{A-9}
\end{equation*}
$$

To cast $I_{2}$ into a form suitable for numerical integration on a computer we subtract the singularity at $\theta=\theta^{\prime}$ and perform the remaining integration explicitly, i.e.,

$$
\begin{equation*}
I_{2}=\int_{0}^{\pi}\left[\frac{\cos \psi e^{i k R} 10}{(1-B \cos \psi)^{\frac{1}{2}}}-\frac{1}{\left(1-B+(B / 2) \psi^{2}\right)^{\frac{1}{2}}}\right] d \psi+\int_{0}^{\pi} \frac{d \psi}{\left(1-B+(B / 2) \psi^{2}\right)^{\frac{1}{2}}} \tag{A-10}
\end{equation*}
$$

and

$$
I_{3}=\int_{0}^{\pi} \frac{d \psi}{\left(1-B+(B / 2) \psi^{2}\right)^{\frac{1}{2}}}=\left(\frac{2}{B}\right)^{\frac{1}{2}}\left\{\ln \left[\pi\left(\frac{B}{2}\right)^{\frac{1 / 2}{2}}+\left(\frac{B \pi^{2}}{2}+1-B\right)^{\frac{1}{2}}\right]-\frac{1}{2} \ln (1-B)\right\}(A-11)
$$

That last term in ( $\mathrm{A}-10$ ) is singular and to perform the $\theta$ ' integration we have to subtract this singularity out. Following a similar proceciure as before we finally arrive at the integral equation

$$
\begin{align*}
& \frac{1}{8 \pi} K_{\theta}(\theta)\left\{4 \pi+\int_{0}^{\pi}\left\{\ln \left[1-\cos \left(\theta-\theta^{\prime}\right)\right]-\ln \frac{\left(\theta-\theta^{\prime}\right)^{2}}{2}\right\} \mathrm{d} \theta^{\prime}\right. \\
& -\int_{0}^{\pi}\left[\ln \left(1-\cos \theta \cos \theta^{\prime}\right)-\ln \frac{\theta^{2}+\theta^{\prime} 2}{2}\right] d \theta^{\prime} \\
& \left.-\pi \ln \left(\theta^{2}+\pi^{2}\right)-2 \theta \tan ^{-1} \frac{\pi}{\theta}+2 \theta \ln \theta+2(\pi-\theta) \ln (\pi-\theta)\right\} \\
& +\frac{1}{8 \pi} \int_{0}^{\pi}\left[K_{\theta}\left(\theta^{\prime}\right)-K_{\theta}(\theta)\right] \ln \left(1-B\left(\theta, \theta^{\prime}\right)\right) d \theta^{\prime} \\
& +2 \int_{0}^{\pi} d \theta^{\prime} K_{\theta}\left(\theta^{\prime}\right) \int_{0}^{\pi} d \psi \cos \psi M_{2}\left(\theta, \theta^{\prime}, \psi\right)=i J_{1}(k b \sin \theta) H_{0} \tag{A-12}
\end{align*}
$$

where $M_{2}\left(\theta, \theta^{\prime}, \psi\right)$ has been defined in (A-4). All the integrals involved in (A-12) have nonsingular integrands and they are suitable for numerical integration on a computer.

## Appendix B

1. Retardation time for $\sigma_{B}$

Referring to Fig. 5 we can easily establish the relationships

$$
\begin{equation*}
\psi=23-\pi / 2 \tag{B-1}
\end{equation*}
$$

$$
\begin{equation*}
(1 B) \cos \psi+b \sin \beta=b+2 \delta \tag{B-2}
\end{equation*}
$$

$$
\begin{equation*}
(1 B) \sin \psi=b \cos \beta \tag{B-3}
\end{equation*}
$$

In view of ( $B-1$ ), ( $B-2$ ) and ( $B-3$ ) can be rewritten as

$$
\begin{equation*}
(1 B) \sin 2 \beta+b \sin \beta=b+2 \delta \tag{B-4}
\end{equation*}
$$

$$
\begin{equation*}
-(1 B) \cos 2 \beta=b \cos \beta \tag{B-5}
\end{equation*}
$$

Combining (B-4) and (B-5) we can derive the following relationships

$$
\begin{equation*}
\sin \beta=\frac{1+\sqrt{1+8(1+2 \delta / b)^{2}}}{4(1+2 \delta / b)} \tag{B-6}
\end{equation*}
$$

$$
\begin{equation*}
(1 B) / b=\left[1+(1+2 \delta / b)^{2}-2(1+2 \delta / b) \sin \beta\right]^{\frac{1}{2}} \tag{B-7}
\end{equation*}
$$

Recalling that the zero time for the charge density at $B$ is taken as the moment of arrival of the wavefront at the center of the sphere we understand that the normalized retardation time is

$$
\begin{aligned}
\frac{c T(B)}{b} & =\frac{(1 B)-(12)}{b}=\frac{(1 B)}{b}(1+\cos 2 \beta)=\frac{(1 B)}{b} \cos ^{2} \beta \\
& =\left(1-\frac{\sin \beta}{1+2 \delta / b}\right)\left[1+(1+2 \delta / b)^{2}-2(1+2 \delta / b) \sin B\right]^{1 / 2}
\end{aligned}
$$

The angle $B$ as given by ( $B-6$ ) ranges from $90^{\circ}(\delta / b=0)$ to $45^{\circ}(\delta / b=\infty)$.
2. Retarcation times for $I$ and $\sigma_{T}$

Referring to Fig. 6 we can easily establish the following relationships

$$
\begin{equation*}
\psi=\pi / 2-2 \beta \tag{B-8}
\end{equation*}
$$

$$
\begin{equation*}
b \sin \psi+(23) \cos \psi+b \sin \beta=2 b+2 \delta \tag{B-9}
\end{equation*}
$$

$$
\begin{equation*}
b \cos \psi=(23) \sin \psi+b \cos \beta \tag{B-10}
\end{equation*}
$$

In view of ( $B-8$ ), ( $B-9$ ) and ( $B-10$ ) can be rewritten as

$$
\begin{equation*}
\mathrm{b} \cos 2 \beta+(23) \sin 2 \beta+\mathrm{b} \sin \beta=2 \mathrm{~b}+28 \tag{B-11}
\end{equation*}
$$

$$
\begin{equation*}
b \sin 2 \beta=(23) \cos 2 \beta+b \cos \beta \tag{B-12}
\end{equation*}
$$

Combining ( $B-11$ ) and ( $B-12$ ) we can derive the following relationships

$$
\begin{gather*}
\sin \beta=\frac{1+\sqrt{1+16(1+\delta / b)(1+2 \delta / b)}}{8(1+2 \delta / b)}  \tag{B-13}\\
\frac{(23)}{b}=2\left\{\left[1+(\delta / b)^{2}\right]-(1+\delta / b) \sin \beta\right\}^{\frac{1}{2}} \tag{B-14}
\end{gather*}
$$

Recalling that the zero time for the current is taken as the moment of arrival of the wavefront at point 4 the normalized retardation time is given by

$$
\frac{C T(I)}{b}=\psi+\frac{(12)+(23)}{b}=\psi+(1-\cos \beta)+\frac{(23)}{b}
$$

where $\psi=\pi / 2-2 \beta$, $\beta$ is given by $(B-13)$ and (23)/b by ( $B-14$ ). The angle $\beta$ ranges from $39.821^{\circ}$ for $\delta / b=0$ to $45^{\circ}$ for $\delta / b=\infty$.

The retardation time for point $T$, at the top of the sphere, is

$$
\frac{c T(T)}{b}=\frac{c}{b} T(I)+\pi / 2-1
$$

We have added $\pi / 2$ to take into consideration the additional path of $90^{\circ}$ and subtracted unity to account for the fact that the zero time for the charge density at point $T$ is the instant of arrival of the wavefront at the center of the sphere.

## Appendix C

The specular reflection path is the least time path. Consider an alternate path which deviates from the specular reflection path by an angle $\Delta B$ (Fig. 4). We want to compare (1B) to (13) + (2B). We can easily derive the relationships

$$
\begin{gather*}
(1 B)^{2}=b^{2}+(2 \delta+b)^{2}-2 b(2 \delta+b) \sin \beta  \tag{C-1}\\
(2 B)^{2}=b^{2}+(2 \delta+b)^{2}-2 b(2 \delta+b) \sin (\beta+\Delta \beta) \tag{C-2}
\end{gather*}
$$

We expand $\sin (\beta+\Delta \beta)$ and we consider terms up to order $(\Delta \beta)^{2}$

$$
\begin{align*}
\sin (\beta+\Delta \beta) & =\sin \beta \cos \Delta \beta+\cos \beta \sin \Delta \beta \\
& =\sin \beta+\Delta \beta \cos \beta-\frac{(\Delta \beta)^{2}}{2} \sin \beta+0\left[(\Delta \beta)^{3}\right] \tag{C-3}
\end{align*}
$$

In view of ( $\mathrm{C}-3$ ) and ( $\mathrm{C}-1$ ) we can rewrite ( $\mathrm{C}-2$ ) as follows

$$
(2 B)^{2}=(1 B)^{2}-2 b(2 \delta+b) \Delta \beta \cos \beta+b(2 \delta+b) \sin \beta(\Delta \beta)^{2}+O\left[(\Delta \beta)^{3}\right]
$$

or

$$
\begin{equation*}
(2 B)=(1 B)\left\{1-\frac{2 b(2 \delta+b)}{(1 B)^{2}} \cos B(\Delta B)+\frac{b(2 \delta+b)}{(1 B)^{2}} \sin B(\Delta B)^{2}+O\left[(\Delta B)^{3}\right]\right\}^{\frac{1}{2}} \tag{C-4}
\end{equation*}
$$

Employing the binomial expansion $(1+\varepsilon)^{\frac{1}{2}}=1+(1 / 2) \varepsilon-(1 / 8) \varepsilon^{2}+\cdots$ we can rewrite ( $C-4$ ) as

$$
\begin{align*}
(2 B)= & (1 B)\left\{1-\frac{b(2 \delta+b)}{(1 B)^{2}} \cos \beta(\Delta B)+\frac{b(2 \delta+b)}{2(2 B)^{2}} \sin \beta(\Delta B)^{2}\right. \\
& \left.-\frac{1}{8}\left[\frac{2 b(2 \delta+b)}{(1 B)^{2}}\right]^{2} \cos ^{2} B(\Delta B)^{2}+O\left[(\Delta B)^{3}\right]\right\} \tag{C-5}
\end{align*}
$$

From Fig. 4 we see that

$$
\begin{gather*}
(12)=2 b \sin \frac{\Delta B}{2} \\
(13)=(12) \cos \psi=2 b \sin \frac{\Delta B}{2} \cos \left(\pi / 2-\beta-\frac{\Delta B}{2}\right) \\
=b \sin \beta(\Delta B)+\frac{b}{2} \cos \beta(\Delta B)^{2}+O\left[(\Delta B)^{3}\right] \tag{C-6}
\end{gather*}
$$

Combining ( $C-5$ ) and ( $C-6$ ) we find

$$
\begin{align*}
(2 B)+(13)-(1 B)= & b(\Delta \beta)\left[\sin \beta-\frac{2 \delta+b}{(1 B)} \cos \beta\right] \\
& +\frac{1}{2} b(\Delta \beta)^{2}\left[\cos \beta+\frac{2 \delta+b}{(1 B)} \sin \beta-\frac{b(2 \delta+b)^{2}}{(1 B)^{2}} \cos ^{2} \beta\right] \\
& +O\left[(\Delta \beta)^{3}\right] \tag{C-7}
\end{align*}
$$

From Fig. 4 we find that

$$
(2 \delta+b) \cos \beta=(4 B)=(1 B) \sin \beta
$$

This shows that the first order term $\Delta \beta$ is zero which is in accordance with the principle of least action. Next we examine the second order term. We notice that

$$
(2 \delta+b) \sin \beta=b+(14)
$$

$$
(2 \delta+b) \cos \beta=(4 B)
$$

and ( $C-7$ ) gives
$(2 B)+(1 B)-(1 B)=\frac{1}{2} b(\Delta B)^{2}\left[\cos \beta+\frac{(14)}{(1 B)}+\frac{b}{(1 B)}-\frac{b}{(1 B)} \frac{(4 B)^{2}}{(1 B)^{2}}\right]+0\left[(\Delta B)^{3}\right]$

Recalling that $B<\pi / 2$, (4B) < (1B), (14) $>0$ we understand that the second order term is positive If we choose a ray with a variation $\Delta \beta$ in the opposite direction we can similarly show that the second order term is positive. Thus, any paths that deviate from the specular reflection path by $\Delta \beta$ correspond to a longer time.


Figure 1. A metallic object within a two-parallel-plate simulator.


Figure 2. A metallic sphere within a two-parallel-plate simulator.


Figure 3. A metallic sphere above a ground plate illuminated by a monochromatic plane wave.


Figure 4. Geometry depicting the specular reflection path and an alternative path. The bottom pole B first senses the presence of the ground plate, or equivalently the image sphere, through the specular reflection path.


Figure 5. Geometry for the computation of the minimum time specular reflection path.


Figure 6. Geometry for the minimun time path involving the interaction of the sphere and the ground plate, or equivalently the image sphere, at the equator points. The path consists of a specular reflection off the image sphere and an arc path along the principal sphere.


FIGURE 7


FIGURE 8


FIGURE 9


FIGURE 10


FIGURE 11


FIGURE 12


FIGURE 13


FIGURE 14


FIGURE 15


FIGURE 16


FIGURE 17


FIGURE 18


FIGURE 19


FIGURE 20


FIGURE 21


FIGURE 22


FIGURE 23


FTGURE 24


FIGURE 25


FIGURE 26


FIGURE 27


FIGURE 28

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